Development of physical models of electron cooling in collisionless plasma thruster plumes

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- Energetic charges are emitted by plasma thrusters (HETs, GITs) which can damage spacecraft parts upon impact
- Their expansion and acceleration ultimately depend on the electric potential in the plume region:
 - □ In the case of slow CEX ions, potential profile completely determines their trajectory and energy (~ $\Delta \phi$)
 - The potential is governed essentially by the plume itself so the determination of the plasma properties and the electric potential profile are coupled
- Understanding the plasma thruster plume expansion into vacuum is of utmost importance to:
 - Accurately characterize the electric potential around the spacecraft with EP
 - Assess the erosion and contamination of satellite surfaces and objects embedded in it



EP2-UC3M group in Madrid has experience modeling plasma plumes with fluid models (EASYPLUME [1]) and PIC codes (EP2-PLUS [2]):

(b) $log_{10}(\tilde{n}), M_0 = 15, \gamma = 1.2$

EASYPLUME density and velocity curves:



✤ HET and GIT plumes expanding into vacuum are near-collisionless ($λ_{coll} ≫ L$). Thus the plasma is not in local thermodynamic equilibrium (LTE): electron response is *global*

Merino, M., Cichocki, F., and Ahedo, E., "Collisionless Plasma thruster plume expansion model," Plasma Sources Sci. Tech., Vol. 24, No. 3, 2015, pp. 035006
 Cichocki, F., Merino, M., Ahedo, E., Hu, Y., and Wang, J., "Fluid vs PIC Modeling of a Plasma Plume Expansion," 34th IEPC, 2015
 Cichocki, F., Merino, M., and Ahedo, E., "Modeling and simulation of EP Plasma Plume Expansion into Vacuum," 50th AIAA/ASME/SAE/ASEE JPC, AIAA-2014-3828 (2014)



- ★ The electric potential profile depends strongly on the electron temperature T_e in the plume ($\Delta \phi \sim T_e/e$), whose determination in a fluid model requires a closure relation of the type $T_e = T_e(n)$.
- Usually, isothermal or polytropic models are used:
 - □ Isothermal: $T_e = \text{const}$, and $e\phi = T_{e0} \ln(n/n_0)$ (Boltzmann)
 - > Reasonably accurate in the near region; gives infinite potential fall downstream (unphysical), requires ∞ energy. Equivalent to $\gamma_e = 1$
 - □ Polytropic: $T_e = T_{e0}(n/n_0)^{\gamma_e-1}$, and $e\phi = (T_e - T_{e0})\gamma_e/(\gamma_e - 1)$
 - "Adjustable" potential fall downstream with γ_e parameter; no model exists for γ_e (experimental fittings can still used). Not theoretically justified in a collisionless plasma.



[4] Merino, M. and Ahedo, E., "Influence of Electron and Ion Thermodynamics on the Magnetic Nozzle Plasma Expansion," IEEE Trans. Plasma Sci., Vol. 43-1, pp. 244–251 (2015)



Laboratory (and in-flight) experiments can be used to advance the understanding of the plasma plume expansion problem

E.g. European plasma plume database (www.electric-propulsion.eu)

* Measurements in HET plumes confirm non-trivial electron cooling with a non-constant γ_e even downstream, where *B* effects are negligible [5]:



[5] Nackles M.R., et al., "Experimental and Numerical Examination of the BHT-200 Hall Thruster Plume," 43rd AIAA/ASME/SAE/ASEE JPC, AIAA 2007-5305 (2007)



Goal of this work

- ★ The goal of this work is to develop simple kinetic of T_e(n) and φ along a expanding collisionless (λ_{coll} ≫ L), unmagnetized (ℓ_e ≥ L) plasma thruster plume
 - The cooling laws derived can then be used to inform fluid and hybrid PIC/fluid models
- This activity is part of the ESA project "Model and experimental validation of spacecraft-thruster interactions (erosion) for electric propulsion thruster plumes." Main tasks include:
 - 1. Detailed review of existing plume models and data by EP2-UC3M and ADS
 - 2. Development of the plume coolingmodel by EP2-UC3M
 - 3. Experimental validation to be done by ICARE-CNRS and KTH at ESA-EPL with a SPT100
 - 4. A simplified version of the derived $T_e(n)$ laws will be implemented in SPIS by ONERA



Comparison with magnetized plasma expansion

- EP2 has already modeled electron cooling of a collisionless plasma expanding in a slowly convergent-divergent magnetic field kinetically:
 - □ Electrons and ions are assumed fully magnetized and tied to the magnetic lines of the magnetic nozzle: distribution function has 3 dof: $f = f(z, v_{\parallel}, v_{\perp})$
 - The distribution function of each species is computed solving its Vlasov equation for a fixed electric potential axial profile
 - Imposing quasineutrality, the method is iterated until convergence in the electric potential in the steady state



[6] M.Martinez-Sanchez, J.Navarro, and E.Ahedo, "Electron cooling and finite potential drop in a magnetized plasma expansion," Phys. Plasmas 22, 053501 (2015)



Comparison with magnetized plasma expansion

- Each particle conserves its mechanical energy ε and the magnetic moment μ (to first order)
 - Expressing the distribution function in these variables allows for fast, semi-analytical solution
- In the divergent part, the magnetic mirror effect accelerates electrons downstream and the electric potential confines electrons upstream
 - □ In the velocity space, there are inaccessible regions (not connected with the upstream) depending on μ and ε
 - □ Electron phase space is divided into
 - > Free electrons (connect $-\infty$ with $+\infty$)
 - > Confined electrons (connected to $-\infty$ but not with $+\infty$)
 - > Empty regions (connected only with $+\infty$)
 - > Doubly-trapped electrons: those that do not connect neither with $-\infty$ or $+\infty$.



Forbidden region/ trapped particles



Comparison with magnetized plasma expansion

- This fragmentation gives rise to non-trivial heat-fluxes and anisotropy
- ♦ Complex cooling behavior is observed →
 - Electrons are initially quasi-isothermal
 - A region of near-constant γ_e cooling follows
 - □ The whole behavior depends on the m_i/m_e ratio.
- ✤ Anisotropy develops downstream too →





Intended modeling approach

- In a collisionless, unmagnetized plasma plume we cannot study the expansion line by line, as in the magnetized case
 - □ The distribution functions have 5 dof: $f = f(z, r, v_z, v_r, v_\theta)$
- We similarly have the following conserved quantities at a kinetic level:
 - □ Particle mechanical energy, *H*
 - Particle angular momentum about the axis, L
 - In a slowly-diverging plume, there exist an adiabatic invariant J associated to the action integral in the r direction which is conserved to first order



* f as a function of z, r, H, L, J and averaging in the r direction gives a fast, semi-analytical solution along the expansion direction



Intended modeling approach

- The phase space after r-averaging has 1 dof more than the magnetized case, but model still tractable numerically
- From [7], for L=0 in a convergent plasma beam, it is seen that *J* plays an analogous role to μ in magnetized plasma plumes:
 - We expect similar behavior with confined, doubly-trapped, and free electrons in the divergent unmagnetized plume
 - Non-trivial cooling mechanisms will also be present
- EASYPLUME will be used to obtain a first iteration of the electric potential and cross-validation of the results
- Results will be fitted to approximate cooling laws as a function of a reduced number of parameters

r Sketch of trajectory of a confined electron Z

[7] M.Martinez-Sanchez, and E.Ahedo, "Magnetic mirror effects on a collisionless plasma in a convergent geometry," Phys. Plasmas 18, 033509 (2011)



Conclusions and comments

- Non-trivial evolution of distribution functions in a collisionless, unmagnetized plasma thruster plume warrants a kinetic approach
 - Understanding electron cooling is crucial for correct ϕ determination
- EP2-UC3M has devised a method to study the problem. A similar approach has already been successfully applied to magnetic nozzles
 - Unmagnetized case is more complex and includes extra degree of freedom, but an analog between adiabatic invariant *J* and magnetic moment μ exists that can be exploited after *r*-averaging
 - Simplified version of the model (fitting laws) to be implemented in SPIS by ONERA
- Experiments within the ESA project aim to validate model results
 - It is essential to understand that vacuum chamber and environmental effects can (will) affect the *global* electron response and modify the cooling behavior: caution must be put to limit or rule out these effects



Thank you!

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